

Classical Mechanics and Statistical/Thermodynamics

August 2008

Possibly Useful Information

Handy Integrals:

$$\int_0^{\infty} x^n e^{-\alpha x} dx = \frac{n!}{\alpha^{n+1}}$$

$$\int_0^{\infty} e^{-\alpha x^2} dx = \frac{1}{2} \sqrt{\frac{\pi}{\alpha}}$$

$$\int_0^{\infty} x e^{-\alpha x^2} dx = \frac{1}{2\alpha}$$

$$\int_0^{\infty} x^2 e^{-\alpha x^2} dx = \frac{1}{4} \sqrt{\frac{\pi}{\alpha^3}}$$

$$\int_{-\infty}^{\infty} e^{i a x - b x^2} dx = \sqrt{\frac{\pi}{b}} e^{-a^2/4b}$$

Geometric Series:

$$\sum_{n=0}^{\infty} x^n = \frac{1}{1-x} \quad \text{for } |x| < 1$$

Stirling's approximation:

$$n! \approx \left(\frac{n}{e}\right)^n \sqrt{2\pi n}$$

Riemann and related functions:

$$\sum_{n=1}^{\infty} \frac{1}{n^p} \equiv \zeta(p)$$

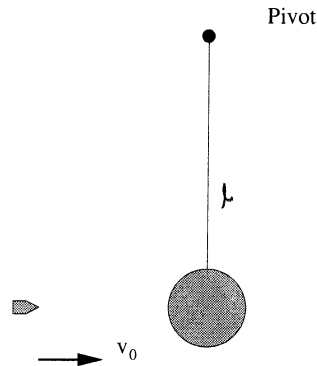
$$\sum_{n=1}^{\infty} \frac{z^p}{n^p} \equiv g_p(z) \quad \sum_{n=1}^{\infty} (-1)^p \frac{z^p}{n^p} \equiv f_p(z)$$

$$g_p(1) = \zeta(p) \quad f_p(1) = \zeta(-p)$$

$\zeta(1) = \infty$	$\zeta(-1) = 0.0833333$
$\zeta(2) = 1.64493$	$\zeta(-2) = 0$
$\zeta(3) = 1.20206$	$\zeta(-3) = 0.0083333$
$\zeta(4) = 1.08232$	$\zeta(-4) = 0$

Classical Mechanics

1. **The ballistic pendulum:** Consider a pendulum with a bob of mass m connected to a frictionless pivot by an ideal massless rigid rod of length ℓ . A projectile of mass ϵm ($0 < \epsilon \ll 1$) moving horizontally at speed v_0 hits the center of the bob, as shown. When it strikes, it becomes imbedded in the bob.

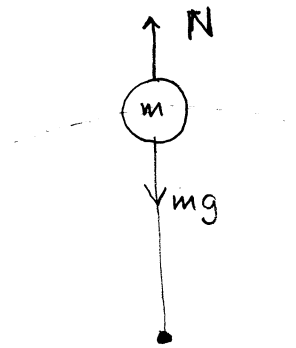


- (a) What is the minimum initial speed of the projectile such that the pendulum will make a full rotation? (2 points)
- (b) The rod is replaced by an ideal massless non-rigid string. What is the minimum initial speed of the projectile such that the pendulum will make a full revolution without the string going slack? (3 points)
- (c) Now assume that projectile rebounds elastically from the bob in the horizontal direction. What is the minimum initial speed of the projectile such that the pendulum will make a full revolution without the string going slack? (2 points)
- (d) Finally, assume that the projectile passes completely through the pendulum bob, in a time $t \ll \sqrt{\ell/g}$. After it exits, it carries with it some of the original mass of the bob, such that the exiting projectile now has a mass $2\epsilon m$ and moves at a speed $3v_0/4$. What is the minimum initial speed of the projectile such that the pendulum will make a full revolution without the string going slack? (3 points)

1.

a) Rigid rod, at the top

$$F_{\text{net}} = 0, \quad v = 0$$



* the bullet sticks inside the bob $\Rightarrow$ inelastic collision

$\rightarrow$ conservation of momentum

$$\epsilon m v_0 = (m + \epsilon m) v$$

$$\Rightarrow v_0 = \left(1 + \frac{1}{\epsilon}\right) v$$

$\rightarrow$ Using conservation of energy on the bob

$$\frac{1}{2} (m + \epsilon m) v^2 = (m + \epsilon m) g 2l$$

$$\Rightarrow v = 2\sqrt{gl}$$

$$\Rightarrow v_0 = 2\sqrt{gl} \left(1 + \frac{1}{\epsilon}\right)$$

b) non-rigid rod

At the top it is moving
in a circular path

$$\text{so, } m'g = \frac{m'v_{\text{top}}^2}{l}$$

$$\Rightarrow v_{\text{top}}^2 = gl$$

Using conservation of momentum

$$v_o = \left(1 + \frac{1}{e}\right) v_{\text{bottom}}$$

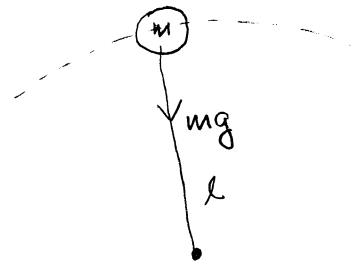
Knowing v_{top} use conservation of energy

$$\frac{1}{2}(m+em)v_{\text{bottom}}^2 = (m+em)g(2l) + \frac{1}{2}(m+em)v_{\text{top}}^2$$

$$\Rightarrow v_{\text{bottom}}^2 = 4gl + gl$$

$$\Rightarrow v_{\text{bottom}} = \sqrt{5gl}$$

$$\therefore v_o = \left(1 + \frac{1}{e}\right) \sqrt{5gl}$$



c) * Elastic collision

$$p_i = p_f$$

$$\Rightarrow \epsilon m v_0 = \epsilon m v_1 + m v_2 \Rightarrow v_0 = v_1 + \frac{1}{\epsilon} v_2$$

$$\Rightarrow v_1 = v_0 - \frac{1}{\epsilon} v_2$$

$$E_i = E_f$$

$$\Rightarrow \frac{1}{2} \epsilon m v_0^2 = \frac{1}{2} \epsilon m v_1^2 + \frac{1}{2} m v_2^2$$

$$\Rightarrow v_0^2 = (v_0 - \frac{1}{\epsilon} v_2)^2 + \frac{1}{\epsilon} v_2^2$$

$$\Rightarrow v_0^2 = v_0^2 + \frac{1}{\epsilon^2} v_2^2 - 2 \frac{v_0 v_2}{\epsilon} + \frac{1}{\epsilon} v_2^2$$

$$\Rightarrow v_0 = \frac{1}{2} \epsilon v_2 + \frac{1}{2} v_2 = \frac{1}{2} v_2 (1 + \epsilon)$$

Now, $v_2 = v_{\text{bottom pendulum}}$

but, $mg = \frac{m v_{\text{top}}^2}{L}$

$$\Rightarrow v_{\text{top}}^2 = g L$$

$$\frac{1}{2} m v_2^2 = \frac{1}{2} m v_{\text{top}}^2 + mg(2L)$$

$$v_2^2 = gL + 4gL = 5gL$$

$$v_0 = \frac{1}{2} (1 + \epsilon) \sqrt{5gL}$$

d) * Momentum conservation

$$\epsilon m v_0 = (m - \epsilon m) v_1 + 2\epsilon m v_2$$

$$\Rightarrow v_0 = \frac{(1 - \epsilon)}{\epsilon} v_1 + 2 v_2$$

$$\text{but } v_2 = \frac{3v_0}{4}$$

$$\Rightarrow v_0 = \left(\frac{1}{\epsilon} - 1\right) v_1 + \frac{3}{2} v_0$$

$$\Rightarrow -\frac{1}{2} v_0 = \left(\frac{1}{\epsilon} - 1\right) v_1$$

$$\Rightarrow v_0 = 2\left(1 - \frac{1}{\epsilon}\right) v_1$$

Now,

$$v_1 = v_{\text{bottom}} = \sqrt{5gl}$$

$$v_0 = 2\left(1 - \frac{1}{\epsilon}\right) \sqrt{5gl}$$

2. The isotropic harmonic oscillator.

- (a) Write the Lagrangian for a point mass m moving under the influence of an isotropic 3-dimensional harmonic oscillator potential

$$V(x, y, z) = \frac{k}{2}(x^2 + y^2 + z^2).$$

There is no external gravitational field. (1 point)

- (b) Using the Lagrange equations of motion show that angular momentum is conserved. i.e.,

$$\frac{d}{dt}\mathbf{L} = \frac{d}{dt}(\mathbf{r} \times m\mathbf{v}) = 0.$$

Because the Lagrangian is invariant under rotations about the origin, you can choose coordinates so that motion is constrained to the x-y plane, i.e., the angular momentum points in the z direction. (3 points)

- (c) For 2-dimensional motion in the x-y plane choose cylindrical polar coordinates and proceed to solve the Lagrange equations of motion. You can leave the solution for $r(t)$ as an integral of the form $t = \int f(r)dr$. (Don't forget to use conservation of energy, E_0 .) (3 points)
- (d) Compute the minimum and maximum values of the radial coordinate r as functions of the constants m, E_0, k, L^z . (3 points)

2.

$$a) \quad L = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - \frac{k}{2} (x^2 + y^2 + z^2)$$

$$x = r \cos \phi \sin \theta$$

$$y = r \sin \phi \sin \theta$$

$$z = r \cos \theta$$

$$\rightarrow \dot{x} = \dot{r} \cos \phi \sin \theta - r \sin \phi \sin \theta \dot{\phi} + r \cos \phi \cos \theta \dot{\theta}$$

$$\rightarrow \dot{y} = \dot{r} \sin \phi \sin \theta + r \cos \phi \sin \theta \dot{\phi} + r \sin \phi \cos \theta \dot{\theta}$$

$$\rightarrow \dot{z} = \dot{r} \cos \theta - r \sin \theta \dot{\theta}$$

$$L = \frac{1}{2} m \dot{r}^2 + \frac{1}{2} m r^2 \dot{\theta}^2 + \frac{1}{2} m r^2 \sin^2 \theta \dot{\phi}^2 - \frac{k}{2} r^2$$

b) trivial

$$c) \quad L = \frac{1}{2} m \dot{r}^2 + \frac{1}{2} m r^2 \dot{\theta}^2 - \frac{k}{2} r^2$$

$$m \ddot{r} - m r \dot{\theta}^2 + \frac{k}{2} = 0$$

$$m r^2 \ddot{\theta} + 2 m r \dot{r} \dot{\theta} = 0$$

$$\Rightarrow r \ddot{\theta} + 2 \dot{r} \dot{\theta} = 0$$

$$\Rightarrow r \frac{d\dot{\theta}}{dr} \dot{r} = -2 \dot{r} \dot{\theta}$$

$$\Rightarrow \frac{d\dot{\theta}}{\dot{\theta}} = -2 \frac{dr}{r}$$

$$\Rightarrow \ln \dot{\theta} = -2 \ln r + \text{const}$$

$$\Rightarrow \dot{\theta} = \frac{1}{r^2} \text{const}$$

$$\leadsto m \ddot{r} - m r \frac{A}{r^4} + \frac{k}{2} = 0$$

$$\Rightarrow m \ddot{r} = + \frac{mA}{r^3} - \frac{k}{2}$$

$$\Rightarrow m \frac{dr}{dr} \dot{r} = \frac{mA}{r^3} - \frac{k}{2}$$

$$\Rightarrow m \dot{r} d\dot{r} = \left(\frac{mA}{r^3} - \frac{k}{2} \right) dr$$

$\Rightarrow$ easier $\Rightarrow$

In cylindrical co-ordinate

$$c) \quad L = \frac{1}{2} m \dot{r}^2 + \frac{1}{2} m r^2 \dot{\theta}^2 + \frac{k}{2} r^2$$

* To evaluate the eqns of motion it is actually easier to work with the Hamiltonian and do Legendre transformation where the Hamiltonian is a const of motion and write it down as a Hamilton-Jacobi eqn

$$H(q, p) = \frac{p_r^2}{2m} + \frac{p_\theta^2}{2mr^2} + \frac{k}{2} r^2$$

$$\theta \text{ cyclic} \quad \dot{p}_\theta = 0 \Rightarrow p_\theta = \alpha_\theta = \text{const.} \quad \downarrow \text{angular momentum}$$

* Use Hamilton's principle func $S(q, \alpha, t)$ as a generating func to get the new Hamiltonian ($K=0$)

* Our H is independent of time

$$\text{so, } S(q, \alpha, t) = \underset{\substack{\uparrow \\ \text{Hamilton's characteristic func}}}{W(q, \alpha)} - E t$$

$$W(q, \alpha) = W_r + \alpha_\theta \theta$$

$$p_r = \frac{\partial W_r}{\partial r}$$

So,

$$\frac{1}{2m} \left(\frac{dW_r}{dr} \right)^2 + \frac{1}{2mr^2} \alpha_\theta^2 + \frac{K}{2} r^2 = E$$

$$\Rightarrow W_r = \int \left(2mE - \frac{\alpha_\theta^2}{r^2} - 2mKr^2 \right)^{1/2} dr$$

$$\Rightarrow R = \frac{\partial W}{\partial P_r} = \frac{\partial}{\partial E} \int ()^{1/2} dr$$

$$\text{but } \dot{R} = \frac{\partial H}{\partial P_r} = \frac{\partial E}{\partial E} = 1$$

$$R = t - t_0$$

$$\Rightarrow t - t_0 = \int \frac{2m}{\sqrt{2mE - \frac{L^2}{r^2} - 2mKr^2}} dr$$

$$(d) \quad E = \frac{1}{2} m \dot{r}^2 + \underbrace{\frac{L^2}{2mr^2} + \frac{K}{2} r^2}_{V_{\text{eff}}}$$

$$\dot{r}^2 = \frac{2E}{m} - \frac{L^2}{m^2 r^2} - \frac{K}{m} r^2 \quad V_{\text{eff}}$$

At the turning pts, $\dot{r} = 0$

$$\Rightarrow r^4 - \frac{2E}{K} r^2 + \frac{L^2}{mK} = 0$$

$$\frac{4L^2}{mK} \times \frac{K^2}{4E^2}$$

$$\Rightarrow r^2 = \frac{(2E/K) \pm \sqrt{(2E/K)^2 - 4L^2/mK}}{2}$$

$$= \frac{E}{K} \left[1 \pm \sqrt{1 - \frac{1}{m} \left(\frac{L}{E} \right)^2} \right]$$

$$\therefore r = \pm \sqrt{\frac{E}{K}} \left[1 \pm \sqrt{1 - \frac{1}{m} \left(\frac{L}{E} \right)^2} \right]^{1/2}$$

$$r_{\text{max}} = \sqrt{\frac{E}{K}} \left[1 + \sqrt{1 - \frac{1}{m} \left(\frac{L}{E} \right)^2} \right]^{1/2}$$

$$r_{\text{min}} = \sqrt{\frac{E}{K}} \left[1 - \sqrt{1 - \frac{1}{m} \left(\frac{L}{E} \right)^2} \right]^{1/2}$$

3. Consider a particle attracted by a fixed gravitating body while also in a uniform gravitational field oriented along the z -axis. The potential energy is of the form:

$$V(r, z) = -m \left(\frac{k}{r} + g z \right)$$

where m is the particle's mass, k and g are constants, and r is the standard radial coordinate:

$$r \equiv \sqrt{x^2 + y^2 + z^2}$$

You are to examine the problem in *cylindrical parabolic coordinates* defined by

$$\zeta \equiv r + z$$

$$\eta \equiv r - z$$

$$\phi \equiv \arctan y/x$$

In these coordinates we may write the Cartesian coordinates as:

$$x = \sqrt{\zeta\eta} \cos \phi$$

$$y = \sqrt{\zeta\eta} \sin \phi$$

$$z = \frac{1}{2}(\zeta - \eta)$$

- (a) Show that the kinetic energy, T , is given by:

$$T = \frac{m}{8} \left[\left(1 + \frac{\zeta}{\eta} \right) \dot{\eta}^2 + \left(1 + \frac{\eta}{\zeta} \right) \dot{\zeta}^2 \right] + \frac{m}{2} \zeta \eta \dot{\phi}^2$$

in these coordinates. (2 points)

- (b) What are the canonical momenta, p_ζ , p_η , and p_ϕ , expressed in cylindrical parabolic coordinates? (2 points)
- (c) Use Hamilton-Jacobi theory to find the constants of the motion. *Hint:* While the total energy E does not separate in these coordinates, $E(\zeta + \eta)$ can be used to produce a quantity that **does** separate. (3 points)
- (d) What is Hamilton's characteristic function associated with ϕ ? (1 point)
- (e) Express Hamilton's characteristic functions associated with ζ , η as definite integrals. (2 points)

3. In Cartesian co-ordinate

$$L = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) + m \left(\frac{k}{\sqrt{x^2 + y^2 + z^2}} + gz \right)$$

In terms of cylindrical parabolic co-ordinate

$$x = \sqrt{\xi \eta} \cos \phi$$

$$y = \sqrt{\xi \eta} \sin \phi$$

$$z = \frac{1}{2} (\xi - \eta)$$

$$r^2 = x^2 + y^2 + z^2 = \xi \eta + \frac{1}{4} (\xi - \eta)^2 = \frac{1}{4} (\xi^2 + \eta^2) + \frac{1}{2} \xi \eta \\ = \frac{1}{4} (\xi + \eta)^2$$

$$\dot{x} = -\sqrt{\xi \eta} \sin \phi \dot{\phi} + \frac{1}{2} \eta (\xi \eta)^{-1/2} \cos \phi \dot{\xi} + \frac{1}{2} \xi (\xi \eta)^{-1/2} \cos \phi \dot{\eta}$$

$$\dot{y} = \sqrt{\xi \eta} \cos \phi \dot{\phi} + \frac{1}{2} \eta (\xi \eta)^{-1/2} \sin \phi \dot{\xi} + \frac{1}{2} \xi (\xi \eta)^{-1/2} \sin \phi \dot{\eta}$$

$$\dot{z} = \frac{1}{2} (\dot{\xi} - \dot{\eta})$$

$$\rightarrow \dot{x}^2 + \dot{y}^2 + \dot{z}^2 = \xi \eta \dot{\phi}^2 + \frac{1}{4} \eta^2 (\xi \eta)^{-1} \dot{\xi}^2 + \frac{1}{4} \dot{\xi}^2 + \frac{1}{4} \xi^2 (\xi \eta)^{-1} \dot{\eta}^2 + \frac{1}{4} \dot{\eta}^2$$

$$= \frac{1}{4} \left(1 + \frac{\xi}{\eta}\right) \dot{\eta}^2 + \frac{1}{4} \left(1 + \frac{\eta}{\xi}\right) \dot{\xi}^2 + \xi \eta \dot{\phi}^2$$

* the mixed terms cancel each other

$$T = \frac{m}{8} \left[\left(1 + \frac{\xi}{\eta}\right) \dot{\eta}^2 + \left(1 + \frac{\eta}{\xi}\right) \dot{\xi}^2 \right] + \frac{m}{2} \xi \eta \dot{\phi}^2$$

* no velocity term in potential

$$b) \quad p_{\xi} = \frac{\partial T}{\partial \dot{\xi}} = \frac{m}{4} \left(1 + \frac{\eta}{\xi}\right) \dot{\xi}$$

$$p_{\eta} = \frac{\partial T}{\partial \dot{\eta}} = \frac{m}{4} \left(1 + \frac{\xi}{\eta}\right) \dot{\eta}$$

$$p_{\phi} = \frac{\partial T}{\partial \dot{\phi}} = m \xi \eta \dot{\phi}$$

* ϕ cyclic

$$\dot{\phi} = 0, \quad p_{\phi} = \alpha_{\phi} = \text{const}$$

$$c) \quad V(r, z) = -m \left(\frac{k}{r} + gz \right)$$

$$= -m \left(\frac{2k}{\xi + \eta} + \frac{g}{2} (\xi - \eta) \right)$$

$$H = \dot{\xi} p_{\xi} + \dot{\eta} p_{\eta} + \dot{\phi} p_{\phi} - L$$

$$= \frac{m}{8} \left(1 + \frac{\xi}{\eta}\right) \dot{\eta}^2 + \frac{m}{8} \left(1 + \frac{\eta}{\xi}\right) \dot{\xi}^2 + \frac{m}{2} \xi \eta \dot{\phi}^2 + \frac{2mK}{(\xi + \eta)} - \frac{mg}{2} (\xi - \eta)$$

$$= \frac{2 p_{\eta}^2}{m \left(1 + \frac{\xi}{\eta}\right)} + \frac{2 p_{\xi}^2}{m \left(1 + \frac{\eta}{\xi}\right)} + \frac{\alpha_{\phi}^2}{2m\xi\eta} + \frac{2mK}{\xi + \eta} - \frac{mg}{2} (\xi - \eta)$$

Using $W(q, \alpha)$ as the generating func

$$p_\eta = \frac{\partial W_\eta}{\partial \eta}, \quad p_g = \frac{\partial W_g}{\partial g}, \quad p_\phi = \alpha_\phi = \frac{\partial W_\phi}{\partial \phi}$$

* In this case, $H = E$

$$E = \frac{2}{m(g+\eta)} \left\{ \eta \left(\frac{\partial W_\eta}{\partial \eta} \right)^2 + g \left(\frac{\partial W_g}{\partial g} \right)^2 + 2m^2 K \right\} \\ + \frac{1}{2m g \eta} \left(\frac{\partial W_\phi}{\partial \phi} \right)^2 - \frac{mg}{2} (g - \eta)$$

$$\Rightarrow \frac{mE}{2} (g+\eta) = \left\{ \right\} + \frac{1}{4} \frac{g+\eta}{g\eta} \left(\frac{\partial W_\phi}{\partial \phi} \right)^2 - \frac{m^2 g}{4} (g^2 - \eta^2)$$

$$\Rightarrow \left[\frac{mE}{2} (g+\eta) - \left\{ \right\} + \frac{m^2 g}{4} (g^2 - \eta^2) \right] \frac{4 g \eta}{g+\eta} = \left(\frac{\partial W_\phi}{\partial \phi} \right)^2$$

$$\frac{\partial W_\phi}{\partial \phi} = \alpha_1$$

$$\left[\frac{mE}{2} (\xi + \eta) - \left\{ \eta \left(\frac{\partial W_\eta}{\partial \eta} \right)^2 + \xi \left(\frac{\partial W_\xi}{\partial \xi} \right)^2 + 2m^2 k \right\} + \frac{m^2 g}{4} (\xi^2 - \eta^2) \right] \frac{4\xi\eta}{\xi + \eta} = \alpha_1^2$$

$$\Rightarrow \frac{mE}{2} \xi - \xi \left(\frac{\partial W_\xi}{\partial \xi} \right)^2 + 2m^2 k + \frac{m^2 g}{4} \xi^2 - \frac{\alpha_1^2}{4\xi}$$

$$= \frac{\alpha_1^2}{4\eta} - \frac{mE}{2} \eta + \eta \left(\frac{\partial W_\eta}{\partial \eta} \right)^2$$

* both sides are independent of each other
and must be equal to a const

$$\leadsto \frac{\alpha_1^2}{4\eta} - \frac{mE}{2} \eta + \eta \left(\frac{\partial W_\eta}{\partial \eta} \right)^2 = \alpha_2$$

$$\leadsto \frac{mE}{2} \xi - \xi \left(\frac{\partial W_\xi}{\partial \xi} \right)^2 + 2m^2 k + \frac{m^2 g}{4} \xi^2 - \frac{\alpha_1^2}{4\xi} = \alpha_2$$

$$d) \quad W_\phi = \alpha_1 \phi$$

$$e) \quad \leadsto \frac{m_E}{2} - \left(\frac{\partial W_g}{\partial g} \right)^2 + \frac{2m_K^2}{g} + \frac{m_g^2}{4} g - \frac{\alpha_1^2}{4g^2} - \frac{\alpha_2}{g} = 0$$

$$\Rightarrow W_g = \int \left(\frac{m_E}{2} + \frac{2m_K^2}{g} + \frac{m_g^2}{4} g - \frac{\alpha_1^2}{4g^2} - \frac{\alpha_2}{g} \right)^{1/2} dg$$

$$\leadsto W_\eta = \int \left(\frac{\alpha_2}{\eta} + \frac{m_E}{2} - \frac{\alpha_1^2}{4\eta^2} \right)^{1/2} d\eta$$

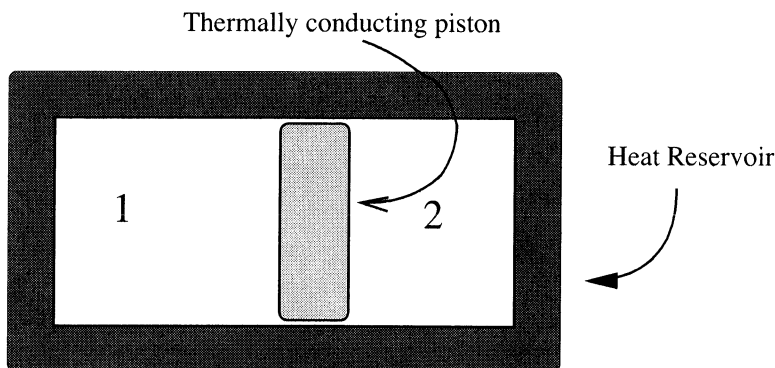
Statistical Mechanics

4. **Helmholtz Free Energy:** The Helmholtz free energy of an ideal monoatomic gas can be written as

$$F(T, V, N) = NkT \left\{ A - \log \left[T^{3/2} \frac{V}{N} \right] \right\}$$

where N is the total number of gas atoms, V is the volume, T is temperature, k is Boltzmann's constant and A is a dimensionless constant.

Consider a piston separating a system into two parts, with equal numbers of particles on the left and the right hand side. The whole system is in good thermal contact with a reservoir at constant temperature T . Initially, $V_1 = 2V_2$. The total volume, $V_{\text{tot}} = V_1 + V_2$, is fixed for this whole problem.



- (a) Calculate the equilibrium position of the piston, once it is released. You must prove your answer, and not simply assert it. (3 points)
- (b) Calculate the maximum available work the system can perform as it changes from the initial condition to the equilibrium position. (3 points)
- (c) Calculate the change in the internal energy, U of gas 1 and gas 2 in the process. (2 points)
- (d) Given your answers above, explain the source of energy for the work done during the expansion. (2 points)

4.

The Helmholtz free energy of an ideal monoatomic gas

$$F(T, V, N) = NKT \left\{ A - \log \left[T^{3/2} \frac{V}{N} \right] \right\}$$

(a)

$$F = E - TS$$

$$\Rightarrow dF = dE - Tds - sdT$$

$$= Tds - PdV + \mu dN - Tds - sdT$$

$$= -PdV + \mu dN - sdT$$

$$\Rightarrow \left(\frac{\partial F}{\partial V} \right)_{N, T} = -P$$

* the system is in thermal equilibrium with the reservoir at temp T

* The system will be in equilibrium if $P_1 = P_2$

$$-P_1 = -NKT \frac{N}{V_1} T^{-3/2} \frac{T^{3/2}}{N} = -\frac{NKT}{V_1}$$

$$\Rightarrow P_1 = \frac{NKT}{V_1} \Rightarrow \text{which is what we expect for an ideal gas}$$

Similarly $P_2 = \frac{NKT}{V_2}$

Thus, at equilibrium, $P_1 = P_2$

$$\Rightarrow V_1 = V_2$$

* Hence the equilibrium pos should be at the middle

(b) * The expansion of V_2 & the compression of V_1 happens isothermally,

$$\begin{aligned} W_1 &= + \int P dV \\ &= + \int_{V_i}^{V_f} \frac{NKT}{V_1} dV \\ &= NKT \ln \left(\frac{V_{1f}}{V_{1i}} \right) \\ &= NKT \ln \left(\frac{3}{4} \right) \end{aligned}$$

$$V_{tot} = V_1 + V_2$$

$\frac{i}{j}$

$$V_{tot} = 3V_2$$

$$V_2 = \frac{V_{tot}}{3}$$

$$V_1 = \frac{2}{3} V_{tot}$$

$\frac{f}{j}$

$$V_{tot} = 2V_2$$

$$V_2 = V_1 = \frac{1}{2} V_{tot}$$

$$\begin{aligned} W_2 &= \int P dV = NKT \ln \left(\frac{V_{2f}}{V_{2i}} \right) \\ &= NKT \ln \left(\frac{3}{2} \right) \end{aligned}$$

$$W_{tot} = NKT \ln \left(\frac{9}{8} \right)$$

(c) $\Delta U = N \frac{f}{2} k \Delta T$ but it is an isothermal process
& $\Delta T = 0$

$$\therefore \Delta U_1 = 0$$

$$\Delta U_2 = 0$$

d) $\Delta U = Q - W$ $Q_1 = NKT \ln\left(\frac{3}{4}\right)$
for $\Delta U = 0$ $Q_2 = NKT \ln\left(\frac{3}{2}\right)$
 $\Rightarrow Q = W$

So, the source of work done is the transfer from the reservoir

5. Consider a gas of N non-interacting **one dimensional** diatomic molecules enclosed in a box of “volume” L (actually, just a length) at temperature T .

- (a) The classical energy for a single molecule is:

$$E(p_1, p_2, x_1, x_2) = \frac{p_1^2}{2m} + \frac{p_2^2}{2m} + \frac{1}{2}K(x_1 - x_2)^2$$

where p_1 and p_2 are the classical momenta of the atoms in one diatomic molecule, x_1 and x_2 are their classical positions, and K is the spring constant. Calculate the specific heat for the gas. (You should assume that $KL^2/2 \gg k_B T$, where k_B is Boltzmann's constant.) (4 points).

- (b) In the quantum limit the energy levels of the molecule are discrete. In a semiclassical approach we can write the energy of one molecule as:

$$E(P, n) = \frac{P^2}{4m} + \hbar\omega\left(n + \frac{1}{2}\right)$$

where P is the momentum of the diatomic molecule (of mass $2m$), and ω is the natural frequency of the oscillator, and n is a non-negative integer ($n \geq 0$). Calculate the specific heat. (4 points).

- (c) Calculate the high and low temperature limits of your result in (b), and explain how they relate to the result of (a). (2 points)

5.

$$a) \quad H = \frac{p_1^2}{2m} + \frac{p_2^2}{2m} + \frac{1}{2} K (x_1 - x_2)^2$$

* Notice: the system has 3 degrees of freedom.
Each atom can move in one direction and
they can vibrate.
1 x 2

Make a co-ordinate transformation such that,

$$X_{cm} = \frac{x_1 + x_2}{2} \quad \text{and} \quad x = x_1 - x_2$$

$$L = \frac{1}{2} (m+m) \dot{X}_{cm}^2 + \frac{1}{2} \left(\frac{mm}{m_1+m_2} \right) \dot{x}^2 - \frac{1}{2} K \dot{x}^2$$

$$= \frac{m}{4} (\dot{x}_1^2 + \dot{x}_2^2 + 2\dot{x}_1\dot{x}_2) + \frac{m}{4} (\dot{x}_1^2 + \dot{x}_2^2 - 2\dot{x}_1\dot{x}_2) - \frac{1}{2} K (x_1 - x_2)^2$$

$$= \frac{1}{2} m (\dot{x}_1^2 + \dot{x}_2^2) + \frac{1}{2} K (x_1 - x_2)^2 \Rightarrow \text{which is the original Lagrangian}$$

Thus, the new Hamiltonian becomes

$$H = \frac{p_{cm}^2}{4m} + \frac{p_x^2}{m} + \frac{1}{2} K x^2$$

$$Z_1 = \frac{1}{2\pi\hbar} \int_0^L \int_{-\infty}^{+\infty} dx_{cm} dp_{cm} e^{-\beta \frac{p_{cm}^2}{4m}} \int_{-\infty}^{+\infty} dx dp_x e^{-\beta \frac{p_x^2}{m}} e^{-\frac{\beta K}{2} x^2}$$

as, $\frac{1}{2} K L^2 \gg K_B T$, we can integrate $x \equiv (x_1 - x_2)$ from $-\infty$ to $+\infty$

$$Z_1 = \frac{L}{2\pi h} \sqrt{4\pi m k_B T} \sqrt{\pi m k_B T} \sqrt{\frac{2\pi k_B T}{m}}$$

$$= \underset{\substack{\uparrow \\ \text{rest of the stuff}}}{\alpha} \beta^{-3/2}$$

$$Z = \frac{1}{N!} Z_1^N = \frac{(\alpha \beta^{-3/2})^N}{N!}$$

$$\Rightarrow E = - \frac{\partial \ln Z}{\partial \beta} = + \frac{3N}{2} \frac{1}{\beta}$$

$$\leadsto C_V = \frac{\partial E}{\partial T} = -k\beta^2 \frac{\partial E}{\partial \beta} = \frac{3Nk}{2}$$

$$b) \quad E(p, n) = \underbrace{\frac{p^2}{4m}}_{H_0} + \underbrace{\hbar\omega(n + \frac{1}{2})}_{H'}$$

$$Z_1^0 = \frac{L}{2\pi\hbar} \sqrt{4\pi m K_B T}$$

$$Z_1' = \sum_{n=0}^{\infty} e^{-\beta\hbar\omega(n + \frac{1}{2})} = e^{-\frac{\beta\hbar\omega}{2}} \sum_{n=0}^{\infty} (e^{-\beta\hbar\omega})^n$$

$$= \frac{e^{-\beta\hbar\omega/2}}{1 - e^{-\beta\hbar\omega}}$$

$$Z_1 = Z_1^0 Z_1' = \underbrace{L \sqrt{\frac{m}{\pi\hbar^2}}}_{\alpha} \beta^{1/2} \frac{e^{-\beta\hbar\omega/2}}{1 - e^{-\beta\hbar\omega}}$$

$$Z^N = Z_1^N = \left(\alpha \beta^{1/2} \frac{1}{e^{\beta\hbar\omega/2} - e^{-\beta\hbar\omega/2}} \right)^N$$

$$E = -\frac{\partial \ln Z}{\partial \beta} = -\frac{\partial}{\partial \beta} \left\{ \ln \beta^{N/2} + \ln (e^{\beta\hbar\omega/2} - e^{-\beta\hbar\omega/2})^{-N} + \ln \alpha \right\}$$

$$= \frac{N}{2} \frac{1}{\beta} + N \frac{\frac{\hbar\omega}{2} (e^{\beta\hbar\omega/2} + e^{-\beta\hbar\omega/2})}{(e^{\beta\hbar\omega/2} - e^{-\beta\hbar\omega/2})}$$

$$= \frac{N}{2} \frac{1}{\beta} + \frac{N\hbar\omega}{2} \coth\left(\frac{\beta\hbar\omega}{2}\right)$$

$$C_V = -k\beta^2 \frac{\partial E}{\partial \beta} = \frac{N}{2} + \frac{N\hbar^2\omega^2}{4} k\beta^2 \operatorname{csch}^2\left(\frac{\beta\hbar\omega}{2}\right)$$

c) for $T \rightarrow \infty$ $\frac{\hbar\omega}{2KT} \ll 1$

$$\cosh x \approx \frac{1}{x} = \frac{2KT}{\hbar\omega}$$

$$C_v \approx -\frac{NK}{2} + \frac{N\hbar^2\omega^2}{4K} \beta^2 \frac{4}{\hbar^2\omega^2\beta^2}$$

$$\approx \frac{NK}{2} + NK$$

$$= \frac{3NK}{2}$$

* As expected in high temp limit the energy levels are close to each other compared to the total "volume" that it can be treated as classical diatomic gas

6. Fermions:

- (a) Show that for any non-interacting spin 1/2 fermionic system with chemical potential μ , the probability of occupying a single particle state with energy $\mu + \delta$ is the same as finding a state vacant at an energy $\mu - \delta$. (2 points)
- (b) Consider non-interacting fermions that come in two types of energy states:

$$E_{\pm}(\vec{k}) = \pm \sqrt{m^2 c^4 + \hbar^2 k^2 c^2}$$

At zero temperature all the states with negative energy (all states with energy $E_{-}(\vec{k})$) are occupied¹ and all positive energy states are empty, and that $\mu(T = 0) = 0$. Show that the result of part (a) above means that the chemical potential must remain at zero for all temperatures if particle number is to be conserved. (2 points)

- (c) Using the results of (a) and (b) above, show that the average excitation energy, the change in the energy of the system from its energy at $T = 0$ in three dimensions is given by:

$$\Delta E \equiv E(T) - E(0) = 4V \int \frac{d\vec{k}}{(2\pi)^3} E_{+}(\vec{k}) \frac{1}{1 + e^{\beta E_{+}(\vec{k})}}$$

(2 points)

- (d) Evaluate the integral above for massless ($m = 0$) particles. (2 points)
- (e) Calculate the heat capacity of such particles. (2 points)

¹Technically this means the total energy of the system diverges. If this bothers you, you can assume some large cut-off to the wavevectors, $\hbar k_{\max} c \gg kT$, which will have no effect on your final answers.

6.

- a) * Probability of finding a single particle state with energy $\mu + \delta$ is

$$P(\mu + \delta) = f(\mu + \delta) = \frac{1}{e^{\beta(E - \mu)} + 1} = \frac{1}{e^{\beta\delta} + 1}$$

- * Probability of finding the state $\mu - \delta$ vacant is

$$P'(\mu - \delta) = 1 - \frac{1}{e^{\beta\delta} + 1} = \frac{e^{-\beta\delta}}{e^{\beta\delta} + 1} = \frac{1}{e^{\beta\delta} + 1}$$

$$\therefore P(\mu + \delta) = P'(\mu - \delta)$$

- b) At, $T = 0$, $\mu = 0$

$$\therefore P(\delta) = P'(-\delta)$$

- * So if N is the the no of particles in E_- state then in order to conserve particle no this is the same amount of particles NOT present in the state E_+

$$N P(E_-) = N (1 - P(E_+)) \quad \text{where, } E_- = -E_+$$

$$\Rightarrow \frac{N}{e^{\beta(E_- - \mu)} + 1} = N \left(1 - \frac{1}{e^{\beta(E_+ - \mu)} + 1} \right)$$

$$\Rightarrow \frac{1}{e^{-\beta(E_+ + \mu)} + 1} = \frac{e^{\beta(E_+ - \mu)}}{e^{\beta(E_+ - \mu)} + 1}$$

$$\Rightarrow 1 + e^{-\beta(E_+ + \mu)} = 1 + e^{-\beta(E_+ - \mu)}$$

$$\Rightarrow E_+ + \mu = E_+ - \mu$$

$$\Rightarrow \mu = 0$$

Hence, $\mu = 0$ for any T

c) the average energy is

$$\langle E \rangle = \frac{\overset{g_s}{\downarrow} 2V}{(2\pi)^3} \int_{-\infty}^{+\infty} d^3K f(E) dE$$

$$= \frac{2V}{(2\pi)^3} \int_{-\infty}^{\infty} d\vec{K} f(E)$$

$$f(E_+) |_{T=0} = \frac{1}{e^{+\infty} + 1} \approx 0$$

$$f(E_-) |_{T=0} = \frac{1}{e^{-\infty} + 1} \approx 1$$

$$E(T) = \frac{2V}{(2\pi)^3} \int d^3K \left[E_+ P(E_+) + \underbrace{E_- P(E_-)}_{= -E_+ (1 - P(E_+))} \right]$$

$$= \frac{4V}{(2\pi)^3} \int d^3K E_+ P(E_+) - \frac{2V}{(2\pi)^3} \int d^3K E_+$$

$$E(0) = \frac{2V}{(2\pi)^3} \int d^3K E_-$$

$$= - \frac{2V}{(2\pi)^3} \int d^3K E_+$$

$$\Delta E \equiv E(T) - E(0)$$

$$= 4V \int \left(\frac{dk}{2\pi} \right)^3 E_+(\vec{k}) \frac{1}{1 + e^{\beta E_+(\vec{k})}}$$

d) For $m=0$

$$E_+ = \hbar k c$$

$$dE_+ = \hbar c dk$$

$$k = \frac{E_+}{\hbar c}$$

$$\Delta E = \frac{4V}{(2\pi)^3} 4\pi \int k^2 dk \frac{E_+(k)}{e^{\beta E_+} + 1}$$

$$= \frac{4V}{(2\pi)^3} 4\pi \frac{1}{(\hbar c)^4} \int_0^\infty \frac{E_+}{e^{\beta E_+} + 1} dE_+$$

$$\beta E_+ = x$$

$$dE_+ = \frac{1}{\beta} dx$$

$$E_+^3 = \frac{x^3}{\beta^3}$$

$$= \frac{16\pi V}{8\pi^3} \frac{1}{(\beta \hbar c)^4} \int_0^\infty \frac{x^3}{e^{x+1}} dx$$

$$\Gamma(4) f_4(1) = \zeta(-4)$$

$$= 0$$

$$(1 - 2^{-s}) s! \zeta(s+1)$$