

# SEARCHING FOR GAMMA-RAYS FROM NARROW-LINE SEYFERT 1 GALAXIES

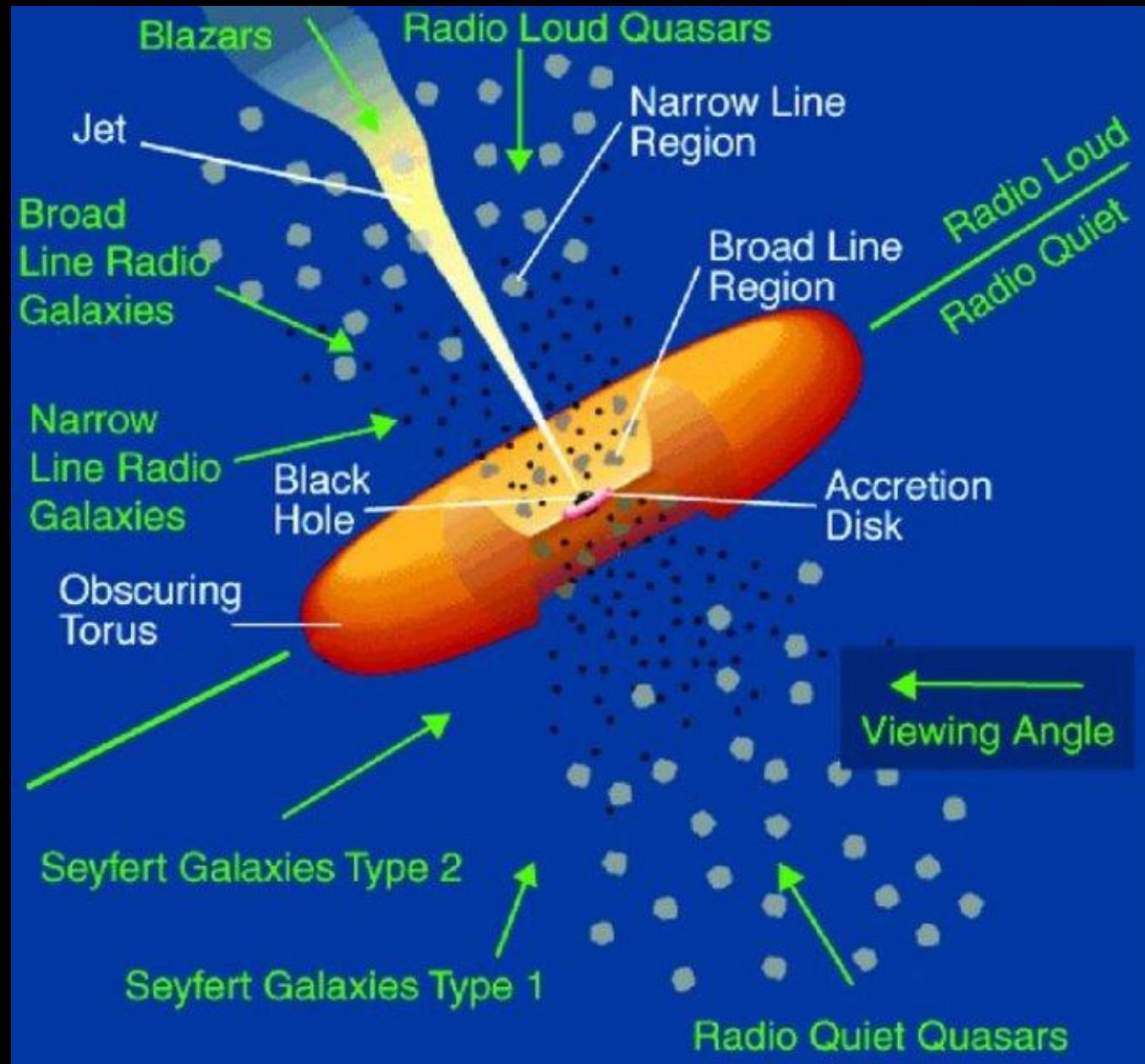
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University of Oklahoma REU

# ACTIVE GALACTIC NUCLEI



Urry C. M. & Padovani P., 1995

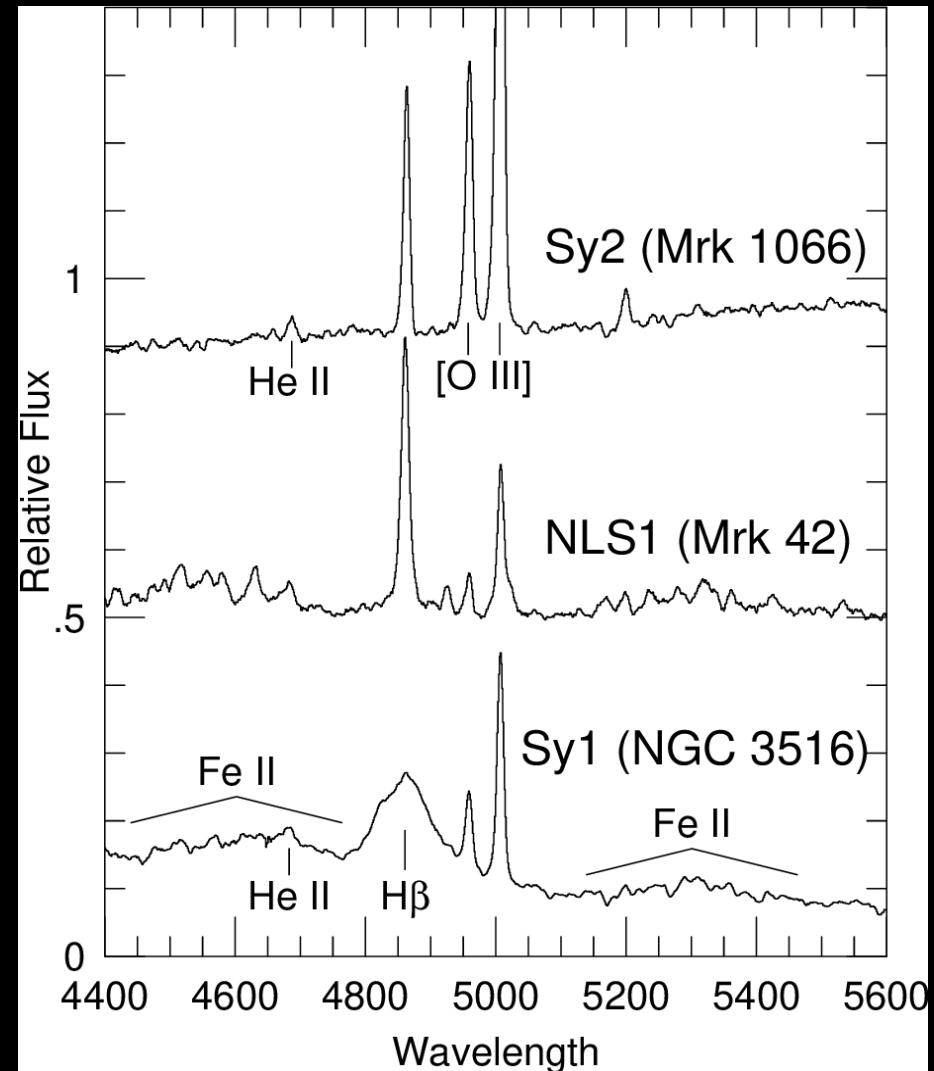
# NLS1

## Classification Criteria

- $FWHM(H\beta) < 2000 \text{ km/s}$
- $F([OIII])/F(H\beta) < 3$
- High  $F([FeII])$

## Characteristics

- SMBH:  $< 10^8 M_\odot$
- Eddington ratio: 0.1-1



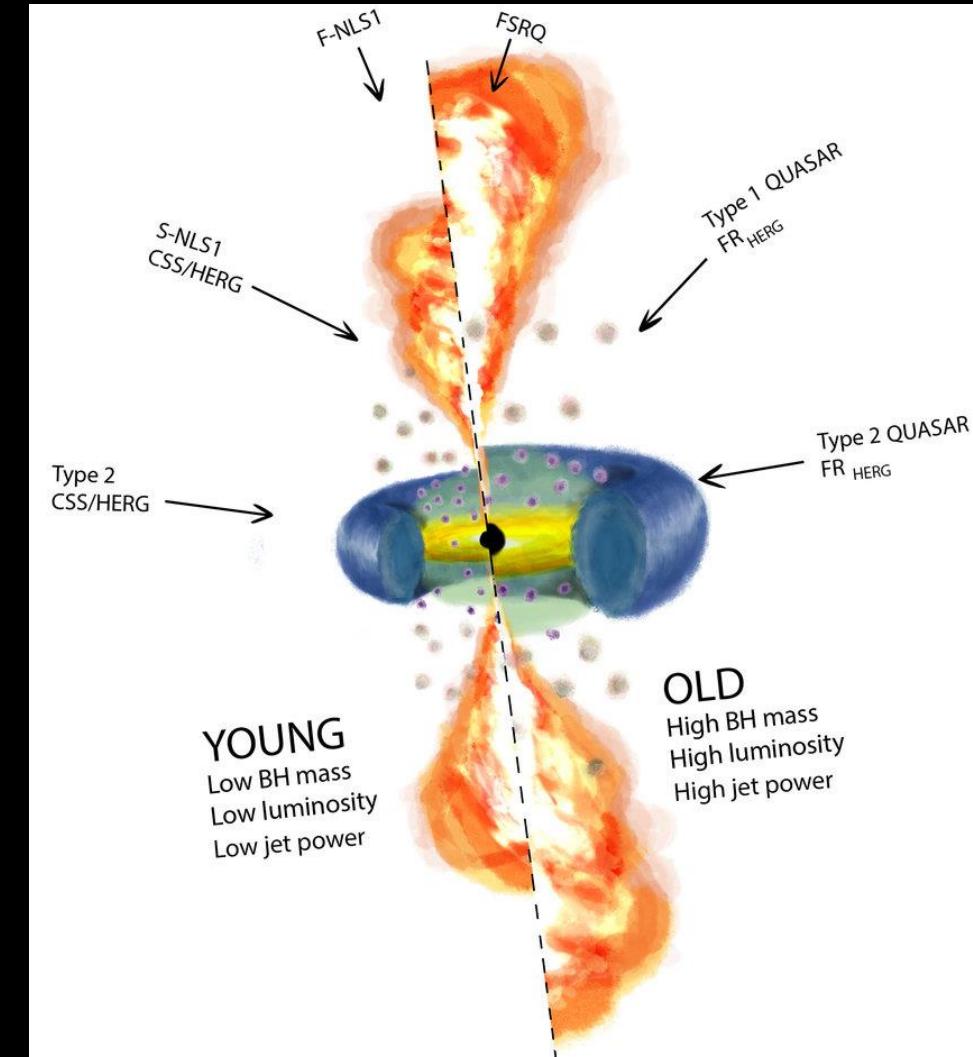
Pogge R. W., 2000

Osterbrock D. E. & Pogge R. W., 1985; Goodrich R. W., 1989;  
Peterson, B. M., et al. 2000; Boroson, T. A. & Green, R. F. 1992

# NLS1

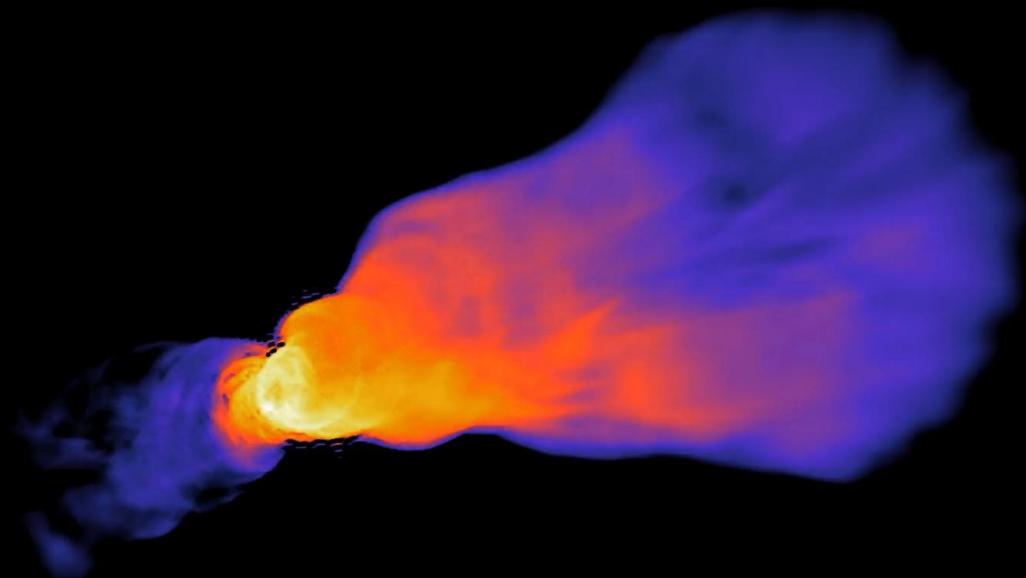
Early evolutionary stage

- Low SMBH mass
- High Accretion rate
- Spiral host galaxies
- Low  $M_{\text{BH}}/M_{\text{bulge}}$



Berton, M., et al. 2017

# LAUNCHING RELATIVISTIC JETS



Andrew Chael. Simulation of M87 SMBH jet.

Blandford-Znajek (BZ)

- Kerr SMBH

Blandford-Payne (BP)

- Accretion disk

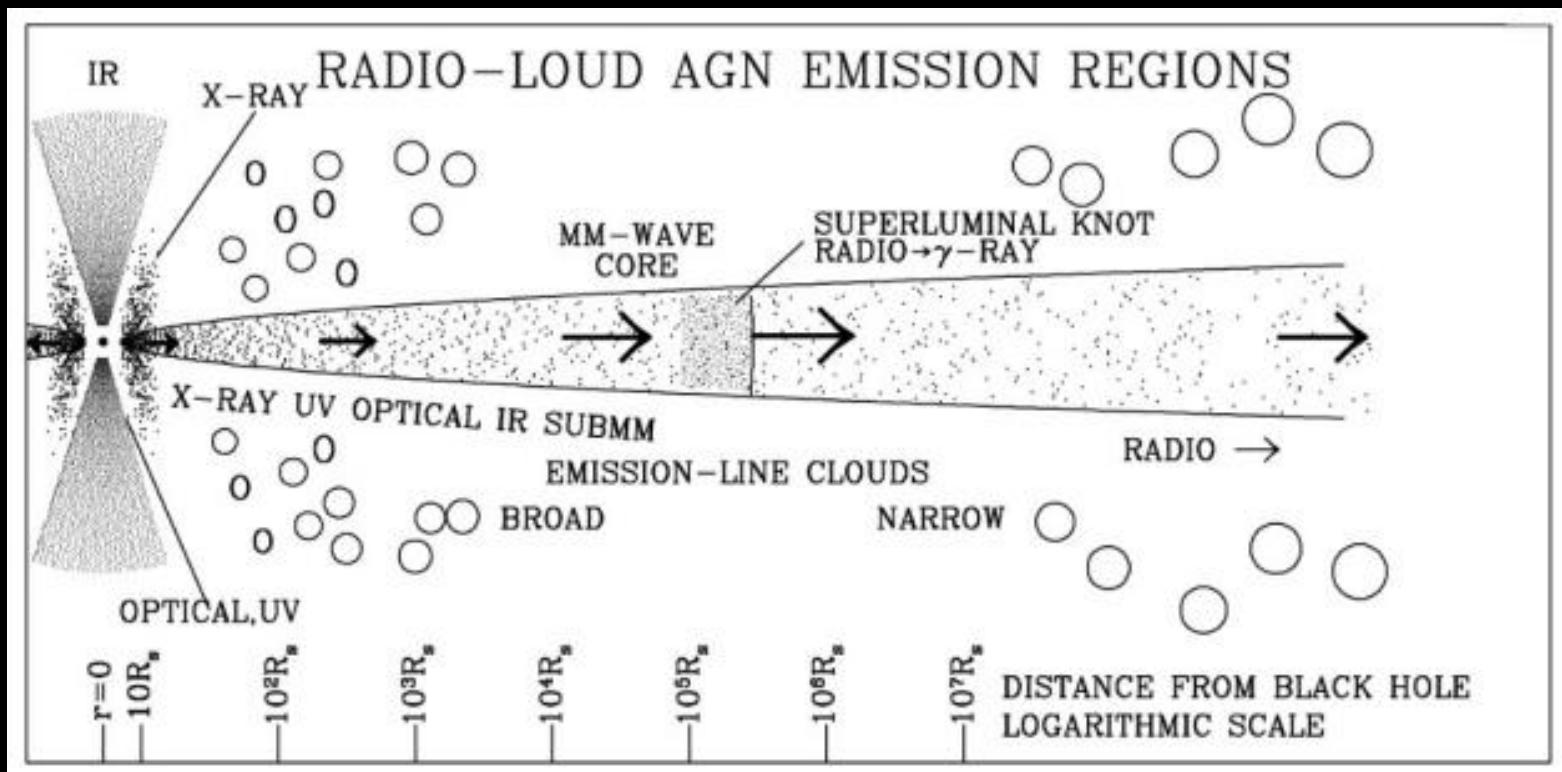
# GAMMA-RAYS

External inverse Compton scattering

- Seed photons from around the jets

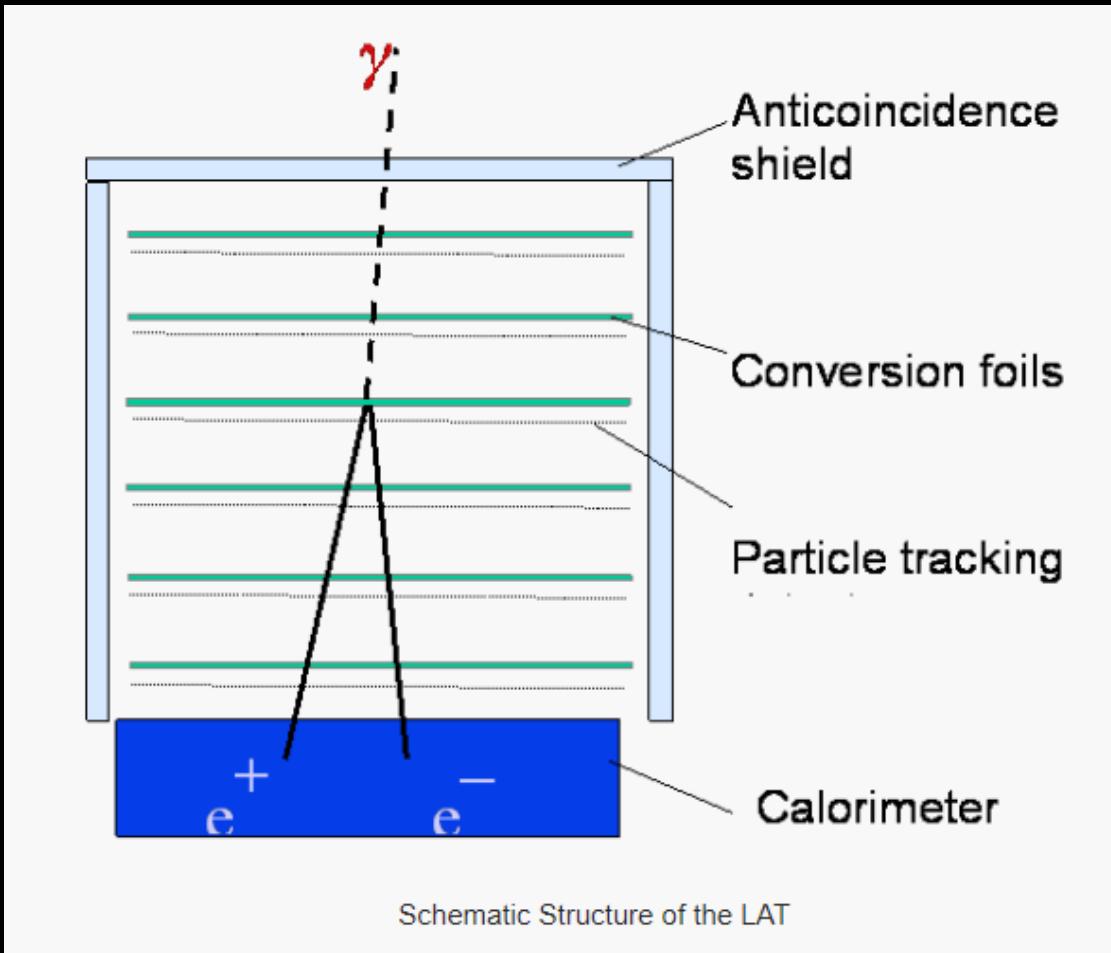
Synchrotron self Compton

- Seed photons from the jets



Marscher, A.P., 2005

# GAMMA-RAY DETECTORS



NASA/DOE/Fermi LAT Collaboration

# FERMI OBSERVATIONS

## Fermi GST

- Launched 2008
- 10 keV to 300 GeV

## Fermi LAT

- 2.4 st FOV
- $0.8 m^2$  effective area



NASA/DOE/Fermi LAT Collaboration

# FERMI OBSERVATIONS

4FGL-DR3

- 14-year catalog
- 1 year+ binning
- Data reduction required for NLS1 detection



# DATA REDUCTION

Maximum Likelihood Estimation

- Maximize likelihood function  $L(\gamma)$
- $TS = -2\log\left(\frac{L(\gamma_0)}{L(\gamma_1)}\right) = 2(l(\gamma_1) - l(\gamma_0))$
- $TS > 25$ : detection ;  $TS > 9$ : candidate

Binning

- 1 year
- 2 months
- 1 week

# THE 7

## Selection Criteria

- Radio silent or radio quiet
- Dense large-scale environments
- Estimated 37GHz flux density or high X-ray/optical

## 2018 candidates/detections:

- SDSS J090113.23 + 465734.7 (TS=10)
- SDSS J122844.81 + 501751.2 (TS=20)
- SDSS J123220.11 + 495721.8 (TS=16)
- SDSS J164100.10 + 345452.7 (TS=39)

# THE 4000

Previous ‘NLS1’ sample

- 11,001
- Highly contaminated

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After cleaning signal-to-noise and remodeling H $\beta$

- 36% remain

# GOALS

Short term

- The 7: NLS1s (mostly) with extreme radio variability

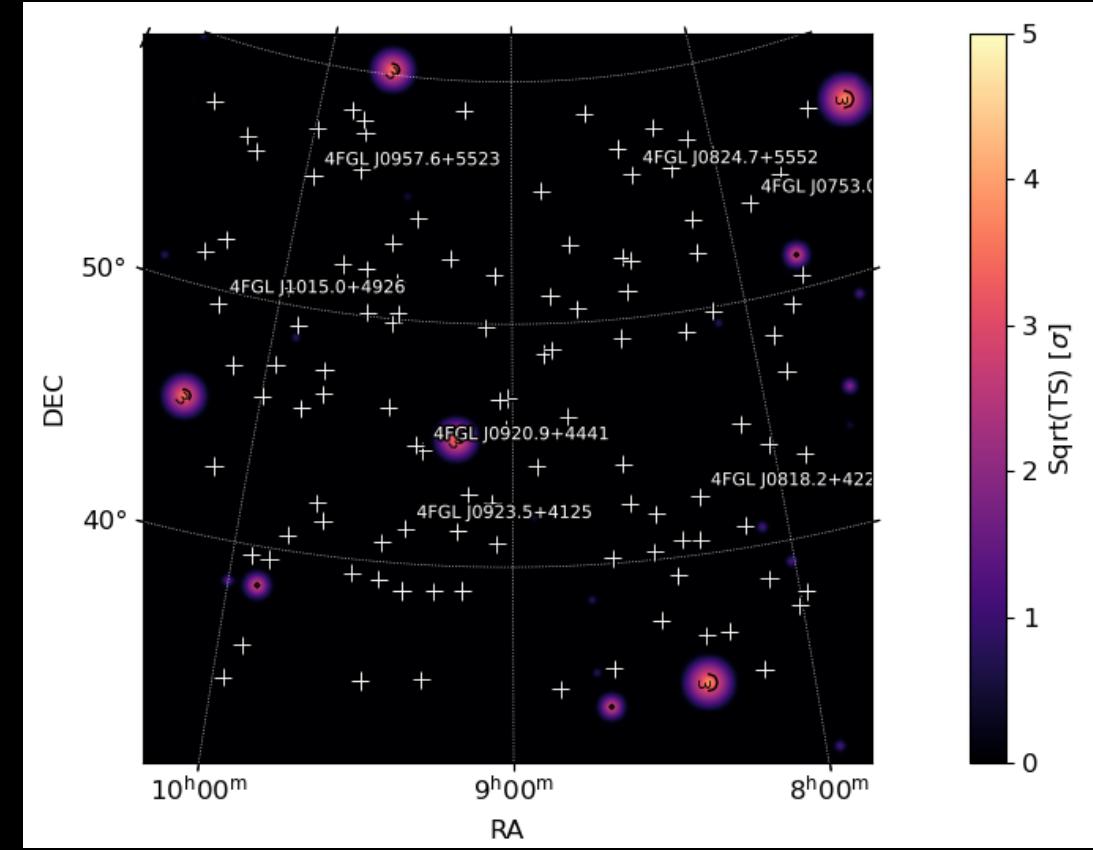
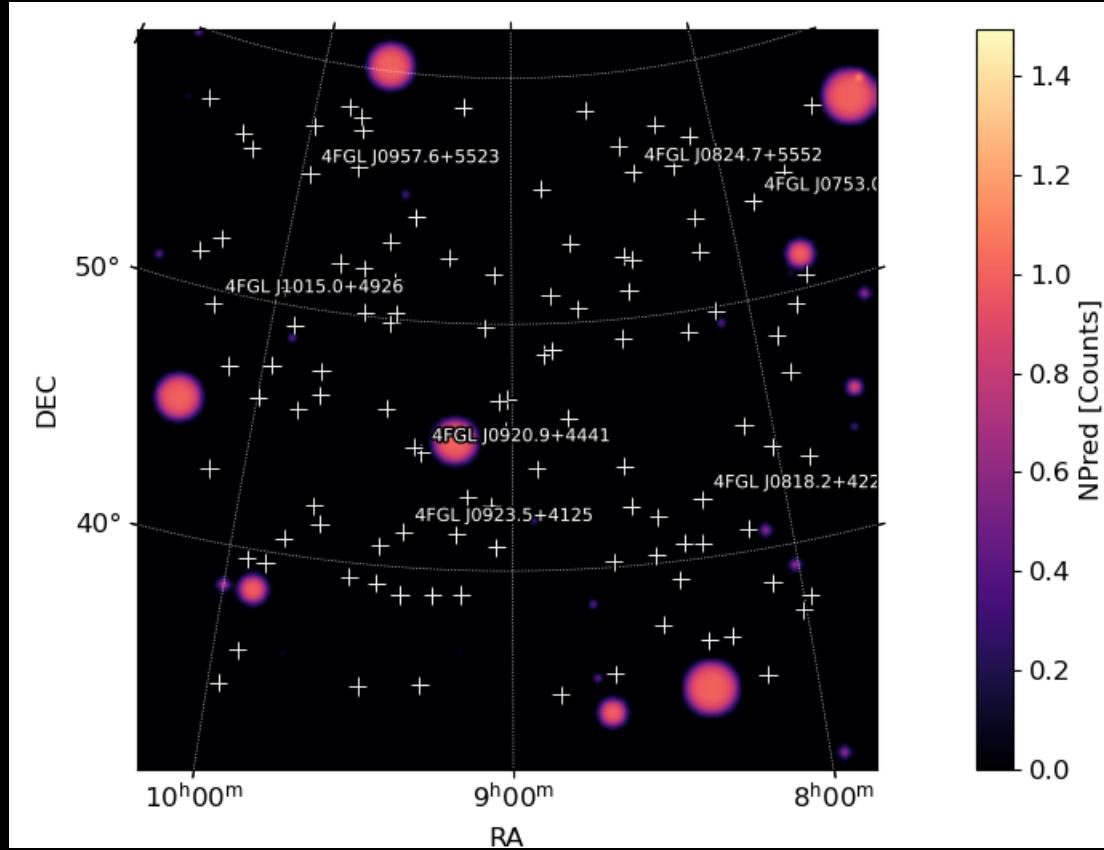
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Long term

- NLS1s that have jets based on radio data
- Largest existing sample of 4000

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# PROGRESS (SO FAR)



# QUESTIONS

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