

Read 6.1-6.3

Finish chapter 5 today

Exam Next Monday (March 11)

7:30 A.M chp 4-5 only

H.W Due today

Next H.W available

(chp 5+6)

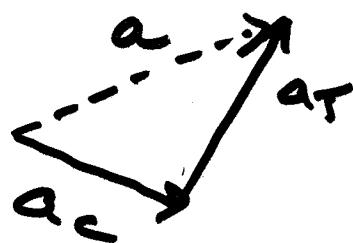
SO FAR we have only looked at uniform circular motion. motion in a circle with constant speed.

NOW look at NON-uniform circular motion motion in a circle with varying speed

For example

car accelerating by pushing on gas pedal while going around a corner

a_T Tangential acceleration



Total acceleration

$$\vec{a} = \vec{a}_c + \vec{a}_T$$

magnitude of a

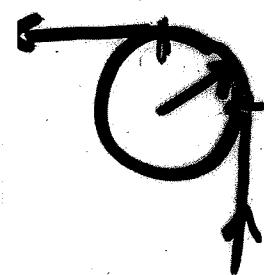
$$a^2 = a_c^2 + a_T^2$$

$$|a| = \sqrt{a_c^2 + a_T^2}$$

ex) while driving your car you accelerate from 0 to 22 m/s along a road that follows a circular path of radius 87 m that turns 90°

a) What is the tangential acceleration? (magnitude)

b) what is the centripetal acceleration (when $v = 15 \text{ m/s}$) (magnitude)



$$v_0 = 0 \text{ m/s} \quad v_f = 22 \text{ m/s} \quad a_t = ? \quad \Delta x = \frac{2\pi r}{4}$$

$$v_f^2 = v_0^2 + 2a_t (\Delta x) \quad \frac{2 \cdot \pi \cdot 87 \text{ m}}{4}$$

$$(22 \text{ m/s})^2 = 0^2 + 2a_t (187 \text{ m}) \quad = 137 \text{ m}$$

$$a_t = 1.7 \text{ m/s}^2$$

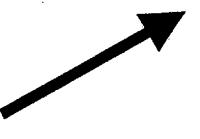
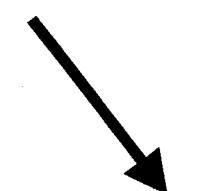
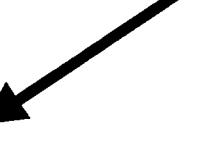
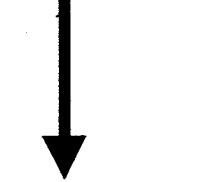
$$\text{b) } a_c = \frac{(15 \text{ m/s})^2}{87 \text{ m}} = 2.58 \text{ m/s}^2 \quad a_c = \frac{v^2}{r}$$

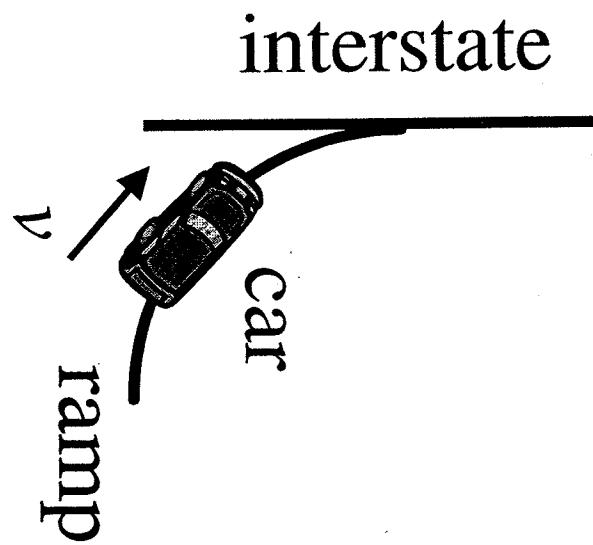
what is acceleration?

$$a = \sqrt{(1.7)^2 + (2.58)^2} = 3.09 \text{ m/s}^2$$

Interactive Question

A car is speeding up as it enters the interstate on a circular entrance ramp as shown in the figure at right. What is the direction of the acceleration of the car when it is at the point indicated?

- A) 
- B) 
- C) 
- D) 
- E) 



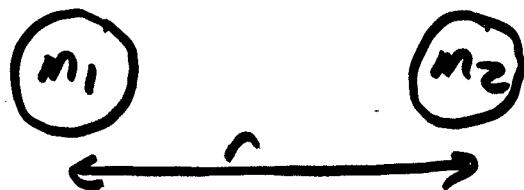
Interactive Question

If an object's speed is increasing, it is possible to have:

- A) Constant tangential acceleration
- B) Constant centripetal acceleration
- C) Neither constant tangential or centripetal acceleration
- D) Both constant tangential and centripetal acceleration

Newton's Law of universal gravitation

$$F = \frac{G m_1 m_2}{r^2}$$



$G = \text{constant}$

$$= 6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2 / \text{kg}^2$$

on surface of earth

$m_p = \text{mass of person}$

A diagram of the Earth represented as an oval with a radius r_E drawn from its center to its surface.

$$F = \frac{G M_E m_p}{(r_E)^2}$$

constant

$$M_E = 5.98 \times 10^{24} \text{ kg}$$

$$r_E = 6.38 \times 10^6 \text{ m}$$

$$\uparrow g$$

$$F = \text{constant} \cdot m$$
$$= \alpha m$$

Circular orbits

what is speed of a satellite in orbit around earth at a distance of 12,200 km above surface?

$$G = 6.67 \times 10^{-11} \text{ N-m}^2/\text{kg}^2$$

$$r_e = 6.38 \times 10^6 \text{ m}$$

$$m_e = 5.97 \times 10^{24} \text{ kg}$$

$$F = \frac{G m_e \cancel{m_s}}{r^2} = \cancel{m_s} \frac{v^2}{r}$$

$$v = \sqrt{\frac{G m_e}{r}}$$

$$v = \sqrt{\frac{6.67 \times 10^{-11} \text{ N-m}^2/\text{kg}^2 \cdot 5.97 \times 10^{24} \text{ kg}}{6.38 \times 10^6 \text{ m} + 1.22 \times 10^7 \text{ m}}}$$

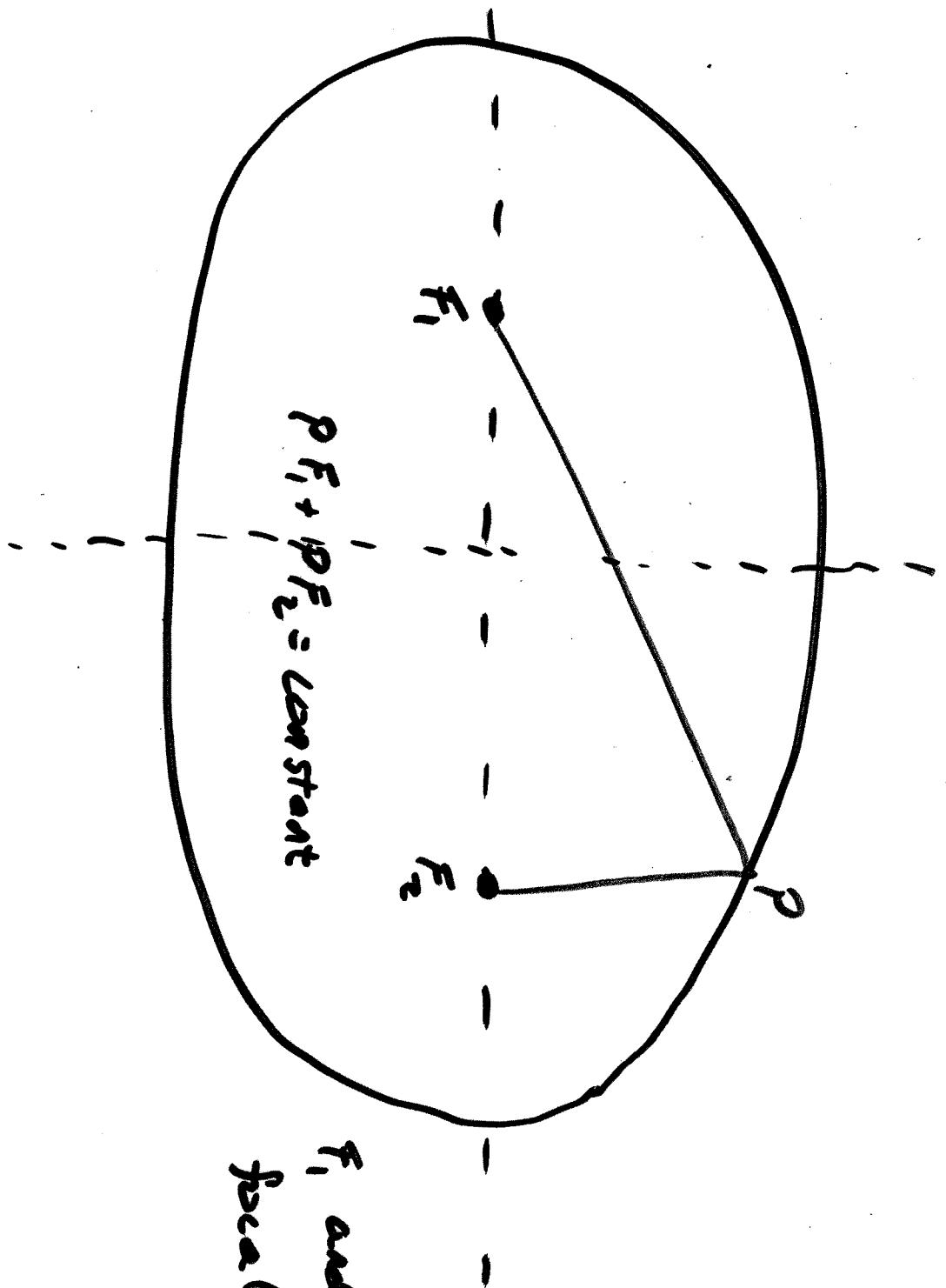
4630 m/s

KEPLER'S 3 LAWS

- 1) Planets move in elliptical orbits with sun at one focal point
- 2) A line drawn from the sun to any planet sweeps out equal areas in equal time intervals
- 3) Square of orbital period of any planet is proportional to the cube of the average distance from the planet to the sun

$$T^2 = \text{constant } r^3 \quad T_1^2 = \text{const } r_1^3$$
$$T_2^2 = \text{const } r_2^3$$

$$\frac{T_1^2}{T_2^2} = \frac{r_1^3}{r_2^3}$$



$$\rho_{F_1} + \rho_{F_2} = \text{constant}$$

F_1 and F_2
focal points

3) is easy to show

Although planets move in elliptical orbits, they are almost circular

$$F = \frac{G m_p m_{\text{sun}}}{r^2} = \frac{m_p V^2}{r}$$

m_p : planet mass

$$V = \sqrt{\frac{G m_{\text{sun}}}{r}}$$

$$V = \frac{2\pi r}{T} = \sqrt{\frac{G m_{\text{sun}}}{r}}$$

square both sides

$$\frac{4\pi^2 r^2}{T^2} = \frac{G m_{\text{sun}}}{r}$$

$$T^2 = r^3 \left(\frac{4\pi^2}{G m_{\text{sun}}} \right)$$

$$\underline{T^2 = \text{const } r^3}$$

Let's see if this works

$$T^2 = \text{const } r^3$$
$$\text{const} = T^2/r^3$$

Mercury: $T = 88$ Days

$$r = 58 \times 10^4 \text{ km}$$

$$\text{const} = \frac{(88)^2}{(58)^3} = \underline{.04}$$

Venus: $T = 226$ days

$$r = 108 \times 10^4 \text{ km}$$

$$\text{const} = \frac{(226)^2}{(108)^3} = \underline{.04}$$

earth $T = 365$ days

$$r = 150 \times 10^4 \text{ km}$$

$$\text{const} = \frac{(365)^2}{(150)^3} = \underline{.04}$$