Classical Mech Main Points

Qualifier

Winter 2012

1 Newtonian Mechanics

- Set $F_N = 0$ to find the point when two objects separate (ex. ball rolls off hemisphere)
- Momentum $(p = mv, L = I\omega)$ is conserved for all collisions; energy is conserved for elastic collisions
- Force $= -\nabla U$
- For periodic motion, if the equation of motion is $\ddot{x} + \xi x = 0$, the frequency is $\omega = \sqrt{\xi}$. If the equation has a term linear in \dot{x} , that is a damping term.

• Power:
$$
P = \frac{dE}{dt} = \frac{\Delta W}{\delta t} = \vec{F} \cdot \vec{v} = \vec{\tau} \cdot \vec{\omega}
$$

1.1 Angular Motion

- Use $v = \omega r$, $x = \theta r$, $a = \alpha r$ for basic angular motion
- Circular motion: $ma = \frac{mv^2}{r} = m\omega^2 r$
- Torque: $\frac{dL}{dt} = \tau = \vec{r} \times \vec{F} = I\alpha = Fd\sin\theta$
- Period $T = \frac{2\pi}{\omega}$
- Remember: it's often easier to find $d \sin \theta$ than to find d and θ separately
- To derive moment of inertia: $I = \int r^2 dm$; solve for dm in terms of dr
- Can still also use $\Sigma F = ma$ if it helps. Consider all forces acting at same point (point particle)
- Orbits: $\frac{\partial^2 V_{eff}}{\partial r^2} > 0$ for stable orbits. Use $\frac{\partial V}{\partial r} = 0$ for circular orbits
- Parallel Axis Theorem: $I_{new} = I_{original} + MR^2$

Helpful moments of inertia:

- sphere: $I = \frac{2}{5}MR^2$
- disc: $I = \frac{1}{2}MR^2$

Rocket Ships: Use $m =$ mass of ship, dm' =ejected mass, v=velocity of ship, $-u=$ ejected mass velocity relative to ship. Then we have:

$$
p_i = p_f \to 0 = (m - dm')(v + dv) + dm'(v - u)
$$
\n(1)

Set $v = 0$ for simplicity, and $dm = -dm'$. After that it's mostly algebra/calculus.

2 Virtual Work

The principle of virtual work presents an alternative to Newtonian solutions for force problems. This method uses the equations:

$$
\delta W = \sum_{i} \vec{F}_{i}^{a} \cdot \delta \vec{r}_{i} = 0 \qquad \delta W = \sum_{i} Q_{i}^{a} \delta q_{i} = 0 \tag{2}
$$

In these equations, \vec{F}_i^a represent the net applied forces, and Q_i^a represent the differentiated constraint equations. Transform the Q_i^a equation into the generalized (simplest) coordinates, and solve the resulting equations.

For example, if the constraint equation is for two blocks connected by a massless rod: $x^2 + y^2 - l^2 = 0$, with $x = l \cos \theta$ and $y = l \sin \theta$:

$$
\delta W = \sum_{i} Q_i^a \delta q_i = 0 \to 2x \delta x + 2y \delta y = 0 \to \delta x \cos \theta + \delta y \sin \theta = 0
$$
\n(3)

2.1 D'Alembert's Principle

The virtual work method given previously works for systems in static equilibrium. To generalize this method to dynamic systems, D'Alembert introduced a new "force of inertia" that modifies the virtual work equation that governs forces:

$$
\delta W = \sum_{i} \left[\vec{F}_{i}^{a} - m_{i} \ddot{r}_{i} \right] \cdot \delta \vec{r}_{i} = 0 \tag{4}
$$

3 Lagrangian & Hamiltonian

3.1 Lagrangian

- $L = T U$
- Euler Lagrange Equation:

$$
\frac{d}{dt}\frac{\partial L}{\partial \dot{q}} - \frac{\partial L}{\partial q} = 0\tag{5}
$$

• We can always add a total time derivative of a function to the Lagrangian for free (without changing equations of motion):

$$
L' = L + \frac{dF(q, \dot{q}, t)}{dt} \tag{6}
$$

This kind of trick can give a simplified Hamiltonian, even making it a constant of the motion.

• A variable q_i is cyclic if it does not appear in the Lagrangian. In that case, the associated momentum p_i is conserved/constant, and subtracting the associated $p_i\dot{q}_i$ transforms the Lagrangian into the Routhian:

$$
p_i = \frac{\partial L}{\partial \dot{q}_i} = \alpha_i \quad \to \quad R = L - \alpha_i \dot{q}_i \tag{7}
$$

3.2 Hamiltonian

- Legendre Transformation: $H = p\dot{q} L$
- \bullet $p_q = \frac{\partial L}{\partial \dot{q}}$
- Hamilton's equations of motion: $\dot{p}_q = -\frac{\partial H}{\partial q}$ and $\dot{q} = \frac{\partial H}{\partial p_q}$
- Solve for q(t) using the E-L equation or Hamilton's equations of motion (take $\frac{dq}{dt}$ and plug in for \dot{p}_q)
- We can see that H is conserved (thus representing the total energy) if $\frac{\partial H}{\partial t} = 0$ and if it includes no terms that depend linearly on a momentum variable (only quadratically).
- We can go farther, and write a momentum-space "Lagrangian", similar to how we did the first Legendre transform: $K(p, \dot{p}, t) = q_i \dot{p}_i + H(q, p, t)$
- $KE_{cylindrical} = \frac{1}{2}m\left(\dot{r}^2 + r^2\dot{\phi}^2 + \dot{z}^2\right)$
- $KE_{spherical} = \frac{1}{2}m\left(\dot{r}^2 + r^2\dot{\theta}^2 + r^2\sin^2\theta\dot{\phi}^2\right)$

3.3 Undetermined Multipliers

If we can't include some constraints when writing the Lagrangian, we have to take these constraints into account in the Euler-Lagrange equation as undetermined multipliers:

$$
\frac{d}{dt}\frac{\partial L}{\partial \dot{q}_i} - \frac{\partial L}{\partial q_i} = Q_i^a + \sum_{j=1}^m \lambda_j a_{ji}
$$
\n(8)

Each λ_j corresponds to each constraint equation f_j , and each a_{ji} corresponds to $\frac{\partial f_j}{\partial q_i}$. Each Q_i^a corresponds to applied forces that cannot be written as part of the potential energy:

$$
Q_i = \frac{\partial \vec{r}_j}{\partial q_i} \cdot \vec{F}_j \tag{9}
$$

A constraint is holonomic if:

$$
\frac{\partial}{\partial y}\frac{\partial f}{\partial x} = \frac{\partial}{\partial x}\frac{\partial f}{\partial y} \tag{10}
$$

3.4 Canonical Transformations

"Guess" the Q & P to transform into in order to make $\frac{\partial H}{\partial t} = 0$. Show canonical by $[Q, P]_{q,p} = 1$

Use existing p and q definitions to find generating functions:

$$
p = \frac{\partial F_1(q, Q)}{\partial q} \qquad P = -\frac{\partial F_1(q, Q)}{\partial P} \tag{11}
$$

$$
p = \frac{\partial F_2(q, P)}{\partial q} \qquad Q = \frac{\partial F_2(q, P)}{\partial P} \tag{12}
$$

$$
q = -\frac{\partial F_3(Q, p)}{\partial p} \qquad P = -\frac{\partial F_3(Q, p)}{\partial Q} \tag{13}
$$

$$
q = -\frac{\partial F_4(p, P)}{\partial q} \qquad Q = \frac{\partial F_4(p, P)}{\partial P} \tag{14}
$$

The generating function(s) result in a new Hamiltonian:

$$
K(Q, P, t) = H(q, p, t) + \frac{\partial F_2}{\partial t}
$$
\n⁽¹⁵⁾

The new Hamiltonian results in corresponding new equations of motion:

$$
\dot{P} = -\frac{\partial K}{\partial Q} \qquad \dot{Q} = \frac{\partial K}{\partial P} \tag{16}
$$

 $H = T + U$ if $\frac{\partial H}{\partial t} = 0$, no explicit time dependence, AND no terms linear in momentum/velocity

3.5 Small Oscillations with Effective Potentials

To find frequency of small oscillations:

- 1. Write the Hamiltonian and find the effective potential, V_{eff} (all terms that depend on q)
- 2. Find $\frac{\partial^2 V_{eff}}{\partial q^2}|_{q=qmin}$ where q represents the variable with small oscillations
- 3. Write the V matrix as:

$$
V = \frac{1}{2}\tilde{V}q^2 = \frac{1}{2}\frac{\partial^2 V_{eff}}{\partial q^2}|_{qmin}q^2
$$
\n(17)

4. Write the T matrix as:

$$
T = \frac{1}{2}\tilde{T}\dot{q}^2\tag{18}
$$

5. Solve for the frequency using \tilde{V} and \tilde{T} :

$$
\tilde{V} - \omega^2 \tilde{T} = 0 \tag{19}
$$

Quick way to get frequency: Make the Lagrangian look like: $L = \frac{1}{2}m'\dot{\eta}^2 - \frac{1}{2}k'\eta^2$. Then $\omega = \sqrt{\frac{k'}{m'}}$

3.6 Variational Calculus

The Euler-Lagrange equation can also solve other physics of path minimization, such as the brachistone problem of minimizing time for a particle in a force field to travel between two points. To use the E-L for this type of problem:

- 1. Write an equation that describes the motion and the element to minimize, such as $dt = \frac{ds}{v}$. The element to minimize should be alone on the LHS.
- 2. Add an integration symbol on both sides: $t = \int \frac{ds}{v}$
- 3. Write the RHS differential in terms of path variables, such as dx and dy , in order to evaluate the integral, such as: $t = \int \frac{\sqrt{1+x'^2}}{\sqrt{2gy}} dy$

4. Use the E-L equation on the integrand, using the appropriate variables, such as: $\frac{\partial F}{\partial x} - \frac{d}{dy} \frac{\partial F}{\partial x'} = 0$

5. Solve the resulting equation by separation of variables, such as $x(y) = \int \sqrt{\frac{y}{(c^2/2g)-y}} dy$

4 Vector Potentials

Remember that the vector potential due to a particle in a magnetic field is:

$$
\vec{A} = -\frac{1}{2}B_0(y\hat{x} - x\hat{y})
$$
\n(20)

And to find the potential, use:

$$
U = q\phi - q\vec{A} \cdot \vec{v} \tag{21}
$$

where ϕ represents the electric potential.

5 Small Oscillations

Standard coordinates define how the blocks are displaced *relative to each other*, while small coordinates (usually η) define how the blocks are displaced *relative to their original equilibrium position*. Start by writing the Lagrangian in standard coordinates, then transform to small coordinates. Then use these notations:

$$
L = \frac{1}{2} \mathbf{T} \dot{\eta}_i \dot{\eta}_j - \frac{1}{2} \mathbf{V} \eta_i \eta_j \tag{22}
$$

Use $\frac{\partial V}{\partial q_i}|_{q0i} = 0$ to find the minimum point q_{0i} , and $\mathbf{V} = \frac{\partial^2 V_{eff}}{\partial q_i^2}$ $\frac{\partial^2 V_{eff}}{\partial q_i^2}\big|_{q0i} = \frac{\partial^2 V_{eff}}{\partial \eta_i \eta_j}$ $\frac{\partial^{\theta} V_{eff}}{\partial \eta_i \eta_j}$ to find **V**.

Then use T and V to solve for the frequency(s):

$$
|\mathbf{V} - \lambda \mathbf{T}| = 0 \tag{23}
$$

where $\lambda = \omega^2$, to solve for the frequencies ω_i . To find the eigenvectors:

$$
(\mathbf{V} - \lambda_i \mathbf{T}) \, \vec{c_i} = 0 \tag{24}
$$

these \vec{c}_i also make up the amplitude ratios for λ_i , $\frac{A_1}{A_2}$.

$$
(\mathbf{V} - \lambda_i \mathbf{T}) \begin{pmatrix} A_1 \\ A_2 \end{pmatrix} = 0
$$
 (25)

To normalize the eigenvectors:

$$
\vec{C}_i = N_i \vec{c}_i \rightarrow \vec{C}_i^T \mathbf{T} \vec{C}_i = 1 \tag{26}
$$

Solve for N_i . Finally, to write the displacement of the system as a function of time:

$$
A_i = \vec{C}_i^T \mathbf{T} \eta(0) \tag{27}
$$

$$
\omega_i^2 > 0 \to \omega_i B_i = \vec{C}_i^T \mathbf{T} \dot{\eta}(0)
$$
\n(28)

$$
\omega_i = 0 \to B_i = \vec{C}_i^T \mathbf{T} \dot{\eta}(0) \tag{29}
$$

The general solution can now be written as:

$$
\vec{\eta}(t) = \sum_{\omega_i^2 > 0} \vec{C}_i (A_i \cos \omega_i t + B_i \sin \omega_i t) + \sum_{\omega_i^2 = 0} \vec{C}_i (A_i + B_i t)
$$
\n(30)

Smaller ω 's correspond to more symmetry in the oscillation mode.

6 Central Forces & the Hamilton Jacobi Equation

Whenever we have two masses exerting a force on each other, we can move into the center of mass reference frame and consider the reduced mass combination acted on by a central force, since the center of mass of the system does not move.

Orbits & Stability

- A circular orbit is stable if $\frac{\partial^2 V_{eff}}{\partial r^2} > 0$
- To find the radius for circular orbit, set $\frac{\partial V}{\partial r} = 0$ and solve for r (can also use Hamilton's equations)
- To find the condition on the radius for circular orbit, find $\frac{\partial^2 V_{eff}}{\partial r^2} > 0$ and substitute in the radius for circular orbit

Steps for Solving Motion with the Hamilton-Jacobi

1. Background: We can transform H without loss of generality to $K = H + \frac{\partial S}{\partial t} = 0$. Assuming then that S, Hamilton's principle/generating function is separable $(S(q,t) = S_1(t) + S_2(q))$ and $p = \frac{\partial S}{\partial q}$, we can rearrange K to be:

$$
\frac{1}{2m} \left(\frac{\partial S_2}{\partial q} \right)^2 + V(q) = -\frac{\partial S_1}{\partial t}
$$
\n(31)

Now the variables are separated, and we can set both sides equal to a constant, E. This makes solving for S_1 and S_2 a matter of maths.

- 2. Write Hamilton's equation, and substitute $\frac{\partial S_2}{\partial q}$ for each p_q term. (S_2 is sometimes referred to as W)
- 3. Separate variables this usually entails writing everything not dependent on r inside a bracket, and setting that bracket equal to α_3 . (This is usually the total angular momentum, which we can see is a constant of the motion by finding $[L, H] = 0$. Or solve so that r is on one side of the equation, and θ and ϕ are on the other side, then set both sides equal to α_3 .
- 4. Assuming W is separable (example $W(r, \theta, \phi) = W_r + W_{\theta} + W_{\phi}$), find integrals defining each component of W.
- 5. Use $p_q = \frac{\partial W}{\partial q}$ to find the meaning of α_2 and α_3 .
- 6. Use the form $\frac{\partial W}{\partial E} = t + \beta$ to solve for the motion of r depending on E and α 's.
- 7. Additional: It may be useful to also remember that $Q = \frac{\partial S_2}{\partial P} = \frac{\partial S_2}{\partial E}$ and $\dot{Q} = \frac{\partial H}{\partial P} = \frac{\partial H}{\partial E}$.

The "action", J is equivalent to $S_2(q)$ as long as $S(q, t)$ is separable:

$$
J = \int pdq = \int PdQ \tag{32}
$$

Given this *J*, the frequency of motion is:

$$
\nu_i = \frac{\partial E}{\partial J_i} \tag{33}
$$

where E came from integrating the action J and solving for $E(J)$.

7 The Poisson Bracket

The poisson bracket is a good method of determining which elements associated with a Hamiltonian are constants of motion:

$$
\frac{du}{dt} = [u, H]_{qi,pi} + \frac{\partial u}{\partial t}
$$
\n(34)

$$
[u, H]_{qi, pi} = \sum_{i}^{n} \left(\frac{\partial u}{\partial q_i} \frac{\partial H}{\partial p_i} - \frac{\partial u}{\partial p_i} \frac{\partial H}{\partial q_i} \right)
$$
(35)

For example, given angular momentum $J = q_1p_2 - q_2p_1$, the poisson bracket of J with H quickly shows that the angular momentum is conserved:

$$
\frac{dJ}{dt} = [J, H]_{qi,pi} = 0
$$
\n(36)

In general, to find whether an element is a constant of motion:

- 1. Write the element A in terms of q_i and p_i
- 2. Write the Hamiltonian according to the physical description
- 3. Find $\frac{dA}{dt} = [A, H]_{qi,pi} + \frac{\partial A}{\partial t}$

For canonical variables:

$$
[q_i, q_j] = 0 \t [q_i, p_j] = \delta_{ij} \t [p_i, p_j] = 0 \t (37)
$$

The poisson bracket also helps verify that transformations are properly canonical:

$$
[Q,P]_{q,p} = 1\tag{38}
$$

8 Extra

8.1 Conservative Forces

A force is conservative if $\vec{\nabla} \times \vec{F} = 0$. In Cartesian coordinates, can find this as: $\frac{\partial F_i}{\partial j} = \frac{\partial F_j}{\partial i}$

8.2 Nonhomogeneous Equations

Solving a non-homogeneous equation requires the combination of a *particular* and a *complementary* solution:

$$
\dot{y} + ay = b \qquad \rightarrow \qquad y(t) = y_p(t) + y_c(t) \tag{39}
$$

- 1. The particular solution should be of the form $y_p(t) = At^2 + Bt + C$, keeping only the terms so that $y_p(t)$ is a polynomial of the same order as the right hand side of the original equation. So in this example, $y_p(t) = C.$
- 2. The complementary solution solves $y(t)$ for the right hand side equalling zero: $\dot{y} + ay = 0$. Solve this the usual way, including the constant of integration.
- 3. Write $y(t) = y_p(t) + y_c(t)$, and substitute these results back into the original equation. Use the original equation and initial conditions to solve for the constants of integration.

Remember that a second derivative equation of motion can be handled as a first derivative equation by writing it in terms of velocity instead of position: $\ddot{y} + a\dot{y} = b \rightarrow \dot{v}_y + av_y = b$

9 Coordinate Systems

9.1 Cartesian

Convert to spherical: $x = r \sin \theta \cos \phi$, $y = r \sin \theta \sin \phi$, $z = r \cos \theta$

Convert to cylindrical: $x = \rho \cos \phi$, $y = \rho \sin \phi$, $z = z$

9.2 Spherical

$$
\hat{r} = \sin \theta \cos \phi \hat{x} + \sin \theta \sin \phi \hat{y} + \cos \theta \hat{z} \tag{40}
$$

$$
\hat{\theta} = \frac{\partial \hat{r}}{\partial \theta} \& \hat{\phi} = \frac{\partial \hat{r}}{\partial \phi} \tag{41}
$$

Derivation of a small chunk of circular area (such as in Kepler's law for orbits):

$$
S = r\theta \to dS = r d\theta \to dA = R^2 d\theta \tag{42}
$$

9.3 Cylindrical

$$
\hat{r} = \cos\theta \hat{x} + \sin\theta \hat{y} \tag{43}
$$

$$
\hat{\theta} = \frac{\partial \hat{r}}{\partial \theta} \tag{44}
$$